

Infrared Atmospheric Sounding Interferometer



IASI



cnes

CENTRE NATIONAL D'ÉTUDES SPATIALES

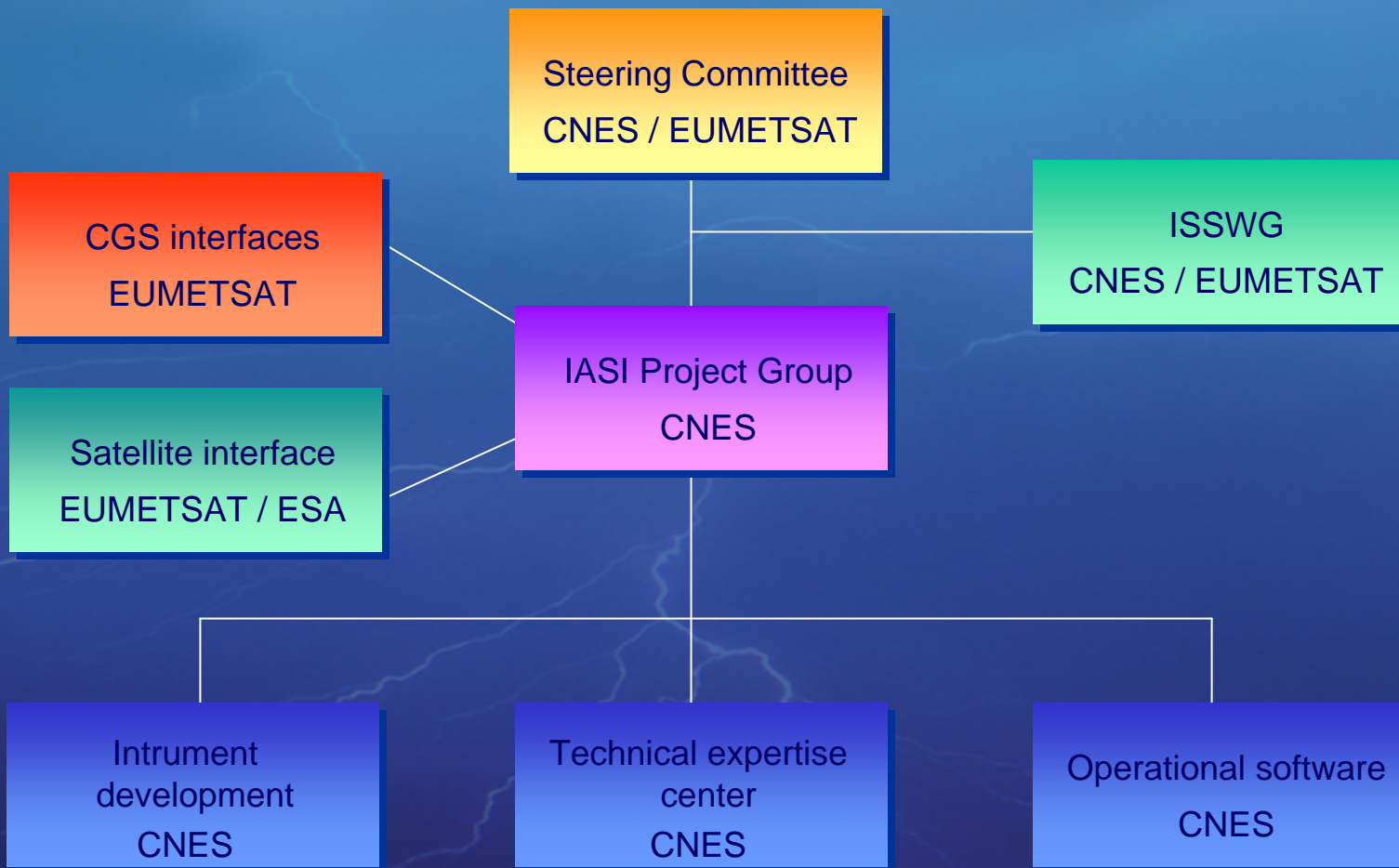


Partners

Prime contractors

Schedule

Partners





Requirements

GEOPHYSICAL VARIABLES	ACCURACY	VERTICAL RESOLUTION	HORIZONTAL SAMPLING
Temperature profile	1K (cloudfree)	1 km	25 km (cloudfree)
Humidity profile	10%(cloudfree)	1-2 km troposphere (cloudfree)	25 km (cloudfree)
Ozone total amount	5%(cloudfree)	N/A	25 km (cloudfree)
Ozone vertical distribution	10%(cloudfree)	2 or 3 pieces of independant information	25 km
Fractional cloud cover	10%		
Cloud top temperature	2K		
Cloud emissivity			
CO, CH ₄ , N ₂ O column	10%	N/A	100 km
SO ₂ , CFCs	10%-20%		
Sea surface temperature	<0.5K (cloudfree)		25 km
Land surface temperature	1K (cloudfree)		25 km
Land surface emissivity	1%		25 km





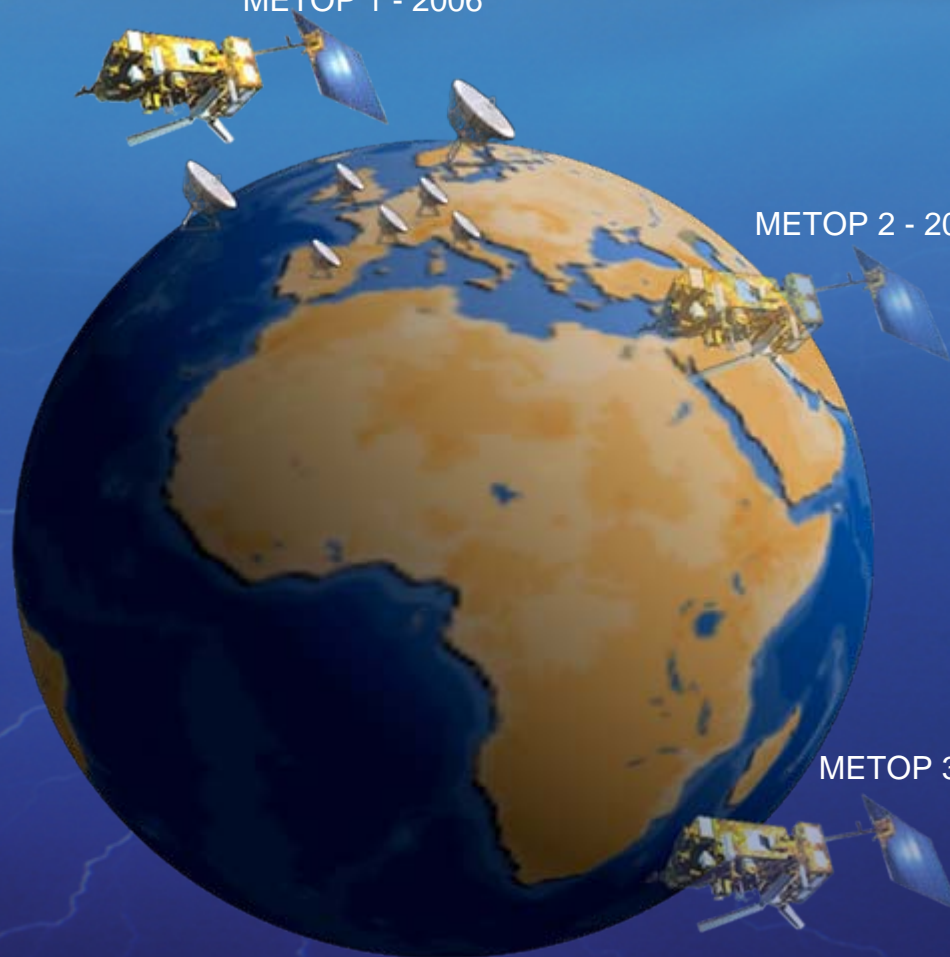
METOP

- Eumetsat Polar System Elements
- 14 years of operation
- >95% reliability on 5 years

METOP 1 - 2006

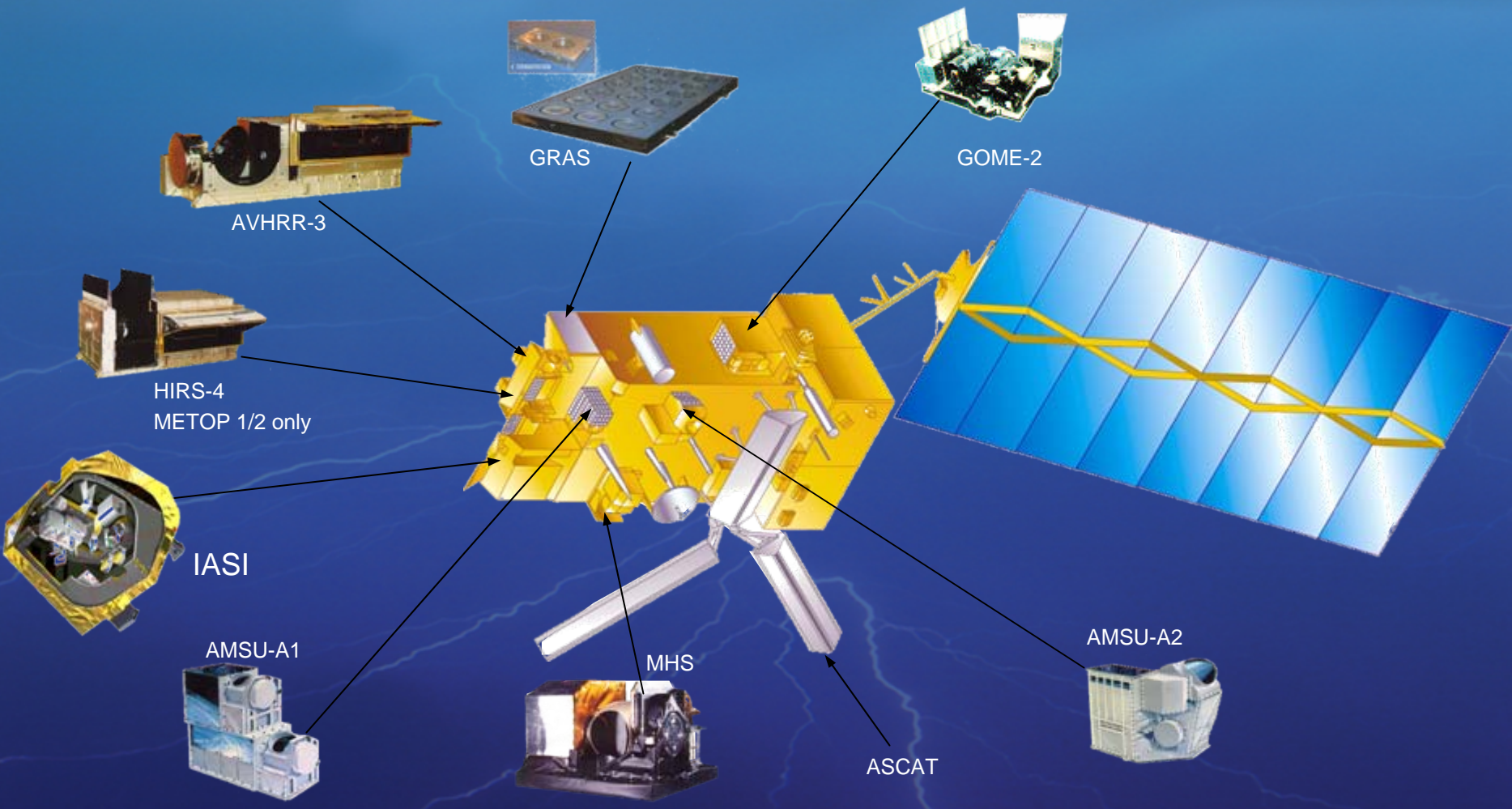
METOP 2 - 2010

METOP 3 - 2015



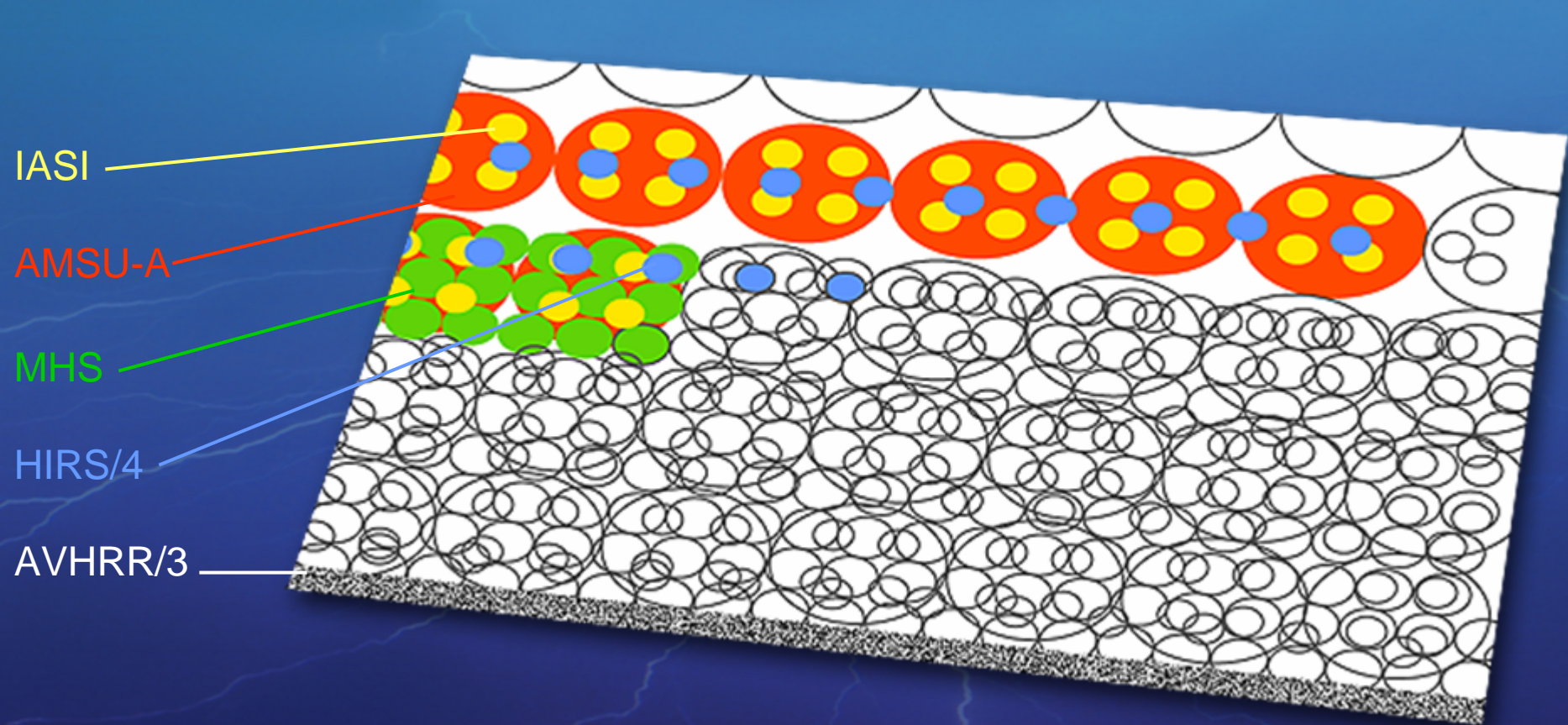


Spacecraft and Instruments





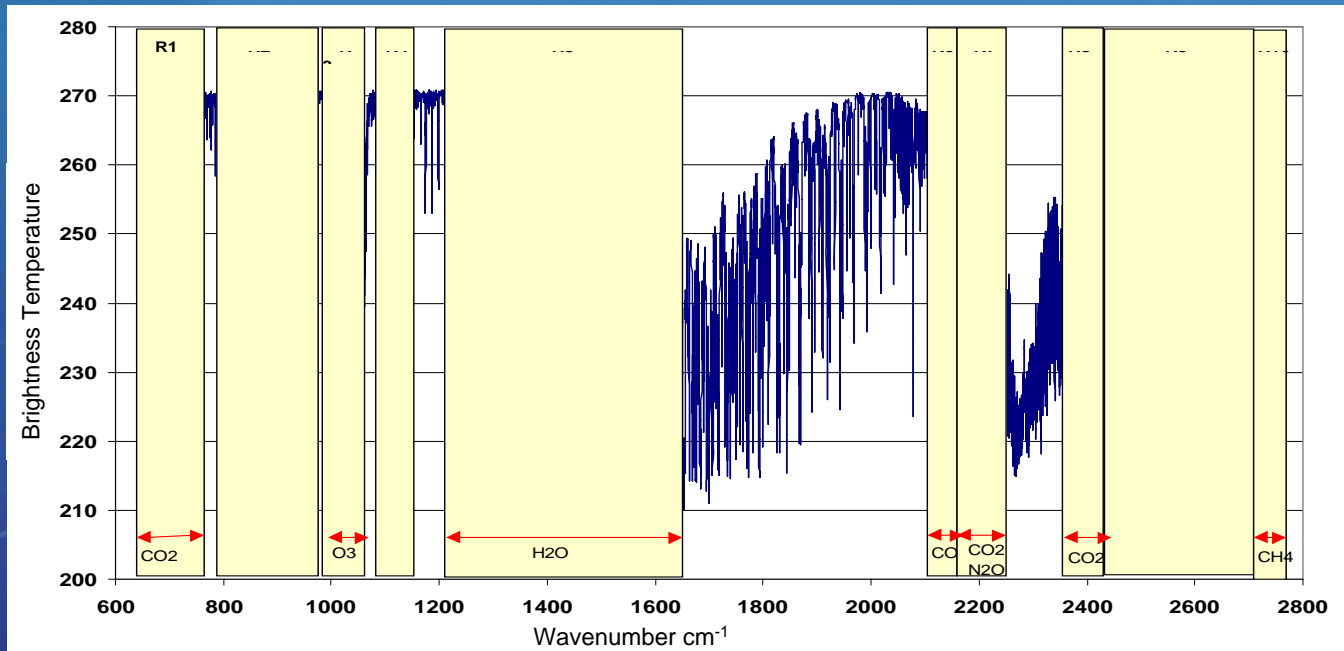
Instrument Scan Patterns





Spectral Range

These products can be derived from high spectral resolution infrared sounder



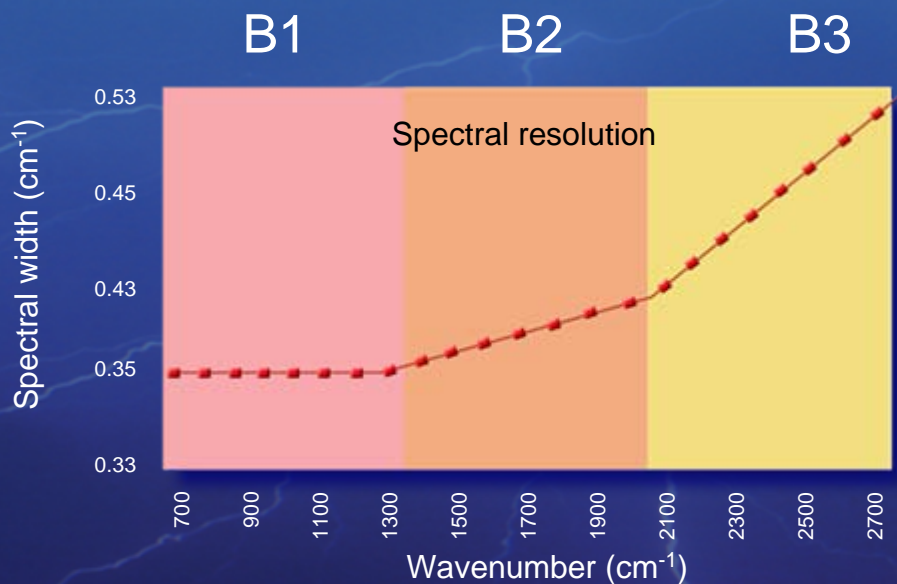
Name	Spectral region	Absorption band	IASI application
R1	650 to 770 cm ⁻¹	CO ₂	Temperature profile
R2	790 to 980 cm ⁻¹	Atmospheric window	Surface and cloud properties
R3	1000 to 1070 cm ⁻¹	O ₃	O ₃ sounding
R4	1080 to 1150 cm ⁻¹	Atmosphere window	Surface and cloud properties
R5	1210 to 1650 cm ⁻¹	H ₂ O	Humidity profile CH ₄ and N ₂ O column amount
R6	2100 to 2150 cm ⁻¹	CO	CO column amount
R7	2150 to 2250 cm ⁻¹	N ₂ O and CO ₂	Temperature profile
R8	2350 to 2420 cm ⁻¹	CO ₂	Temperature profile
R9	2420 to 2700 cm ⁻¹	Atmosphere window	Surface and cloud properties
R10	2700 to 2760 cm ⁻¹	CH ₄	CH ₄ column amount





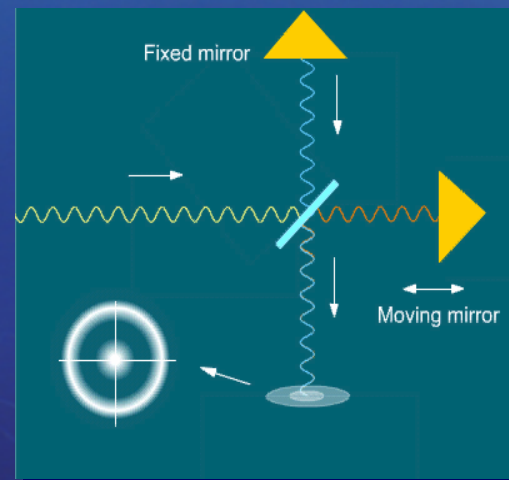
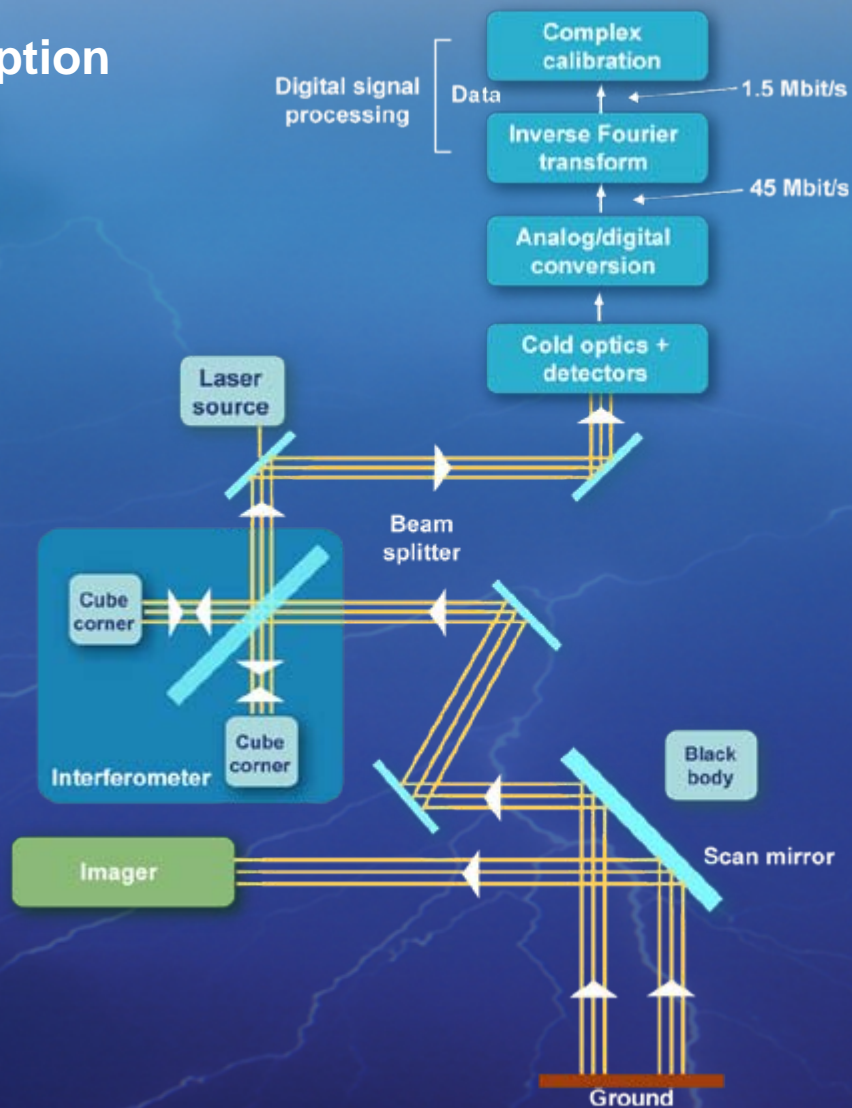
Spectral : resolution (1/2)

- Spectral resolution (FWHM)
 - The spectral resolution specification is based on the line spacing in the 650 cm^{-1} CO_2 absorption band
 - This spacing is about 1.5 cm^{-1}
- Spectral sampling interval = 0.25 cm^{-1}





Functional description





Configuration





Performances Tests

PFM Optical vacuum test : Oct 2003

PFM test on-board Metop: Feb 2004



Courtesy EADS

FM2 optical vacuum test : Sept. 2004



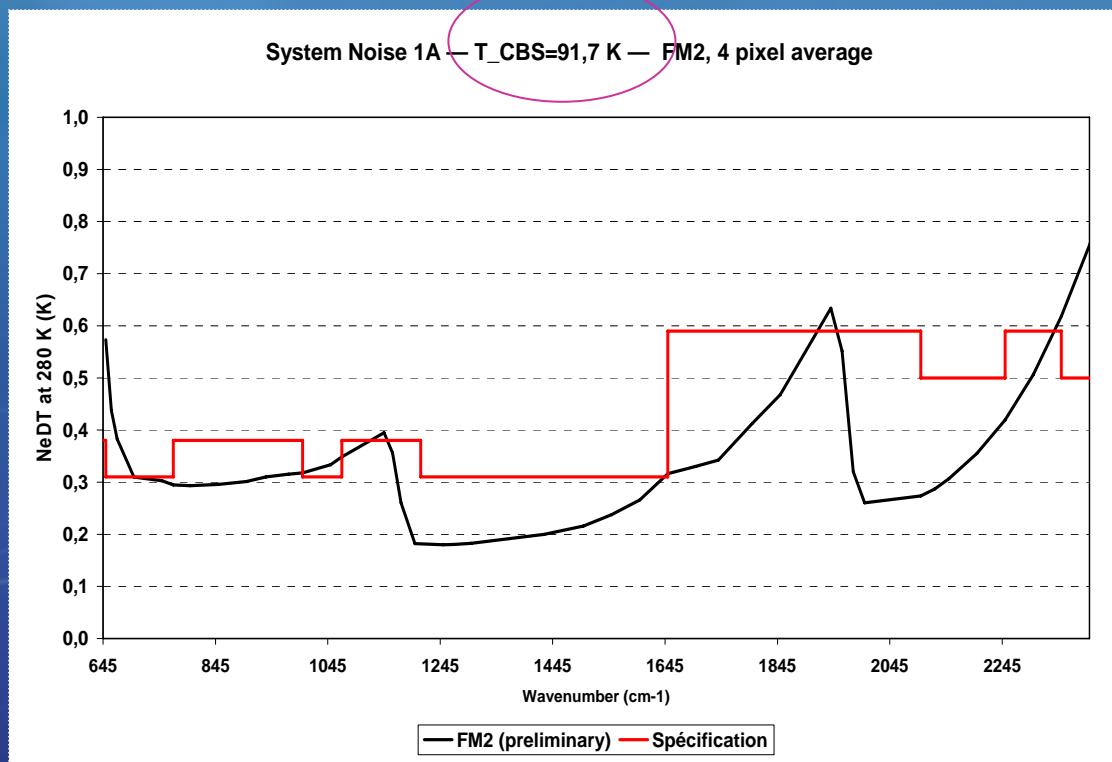
FM3 optical vacuum test : Oct. 2005

Courtesy Alcatel





System radiometric noise for FM2 :



- Raw inteferograms processed with DPS
- Average of 4 pixel *27scan angle*40 lines
- Noise for each line and averaged on 40 lines

After cold box design modifications, ice contamination issue solved





Radiometric calibration

- Interpixel & interscan calibration
- Scan mirror characterization verified
 - External HBB temperature : 85 K
- Impact of the incidence on calibration error
 - Measured for external HBB in SP= 1, 5, 10, 25 positions
 - Incidence : -56.67, -50, -41.67, -16.67 deg
 - External HBB temperature : 294 K & 240 K
- Precautions
 - Measurement repeated twice in order to detect potential thermal drift in the test setup
 - Effect of Earth Panel temperature verified

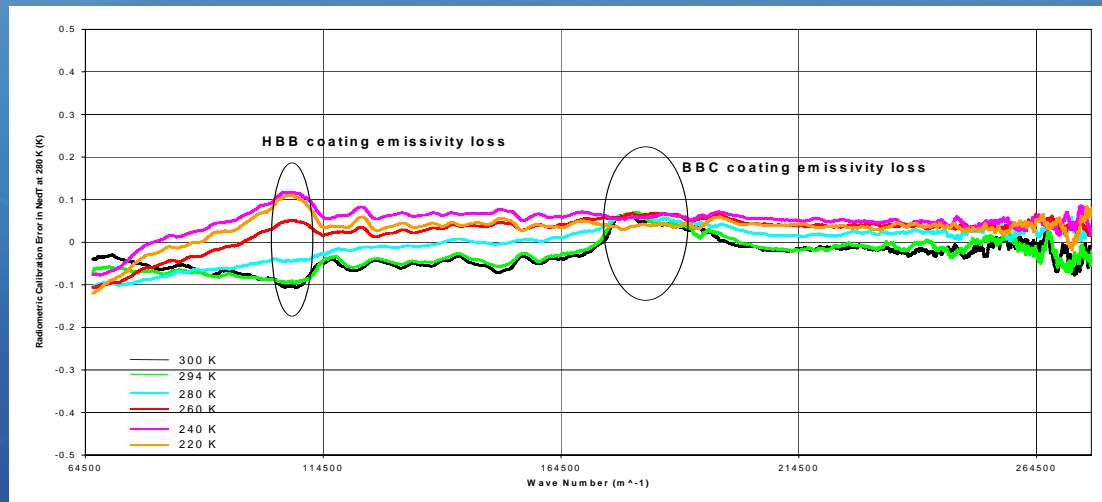




Requirements Principle Design Sub systems Performances Development plan

Radiometric Calibration : Sensitivity to incidence angle, scan mirror reflectivity, or S/I thermal drifts

Raw results for Tscene=220-300k
Pixel 1
Nominal configuration, Cold case

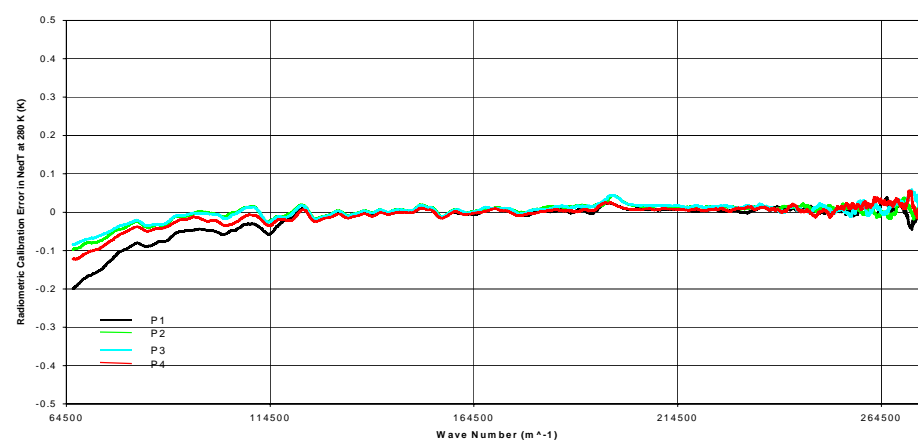
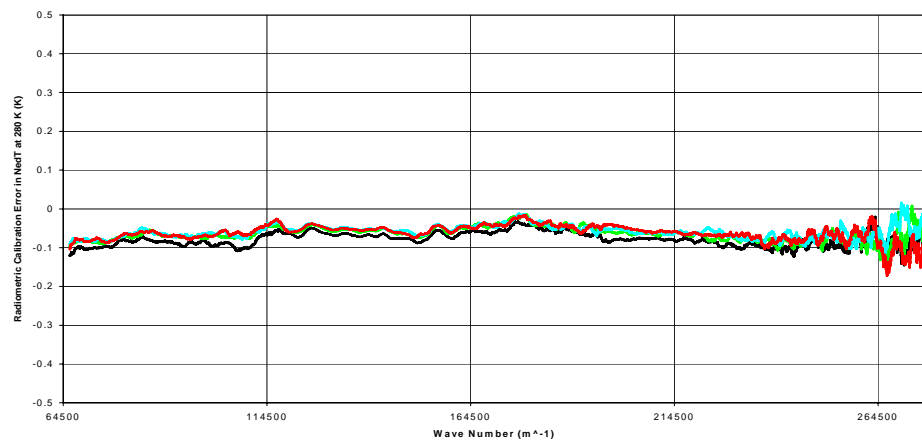


Inter-scan radiometric performance
HBB at 294 K, Pixel 1

(Spec. 0.1 K)

Interpixel radiometric performance
HBB at 294 K, Pixel 1

(Spec. 0.1 K)





Radiometric calibration

- Objectives of the Mission Specification on radiometric calibration are expected to be achieved
- Absolute Calibration error
 - **< 0.5 K**
- Intercalibration errors at a given time for all geophysical conditions
 - **< 0.2 K**
 - **4 pixels, all scan directions, all channels**
- Intercalibration errors over time
 - **< 0.3 K**
 - **Random variations (over orbit period)**
 - **Long term variations**

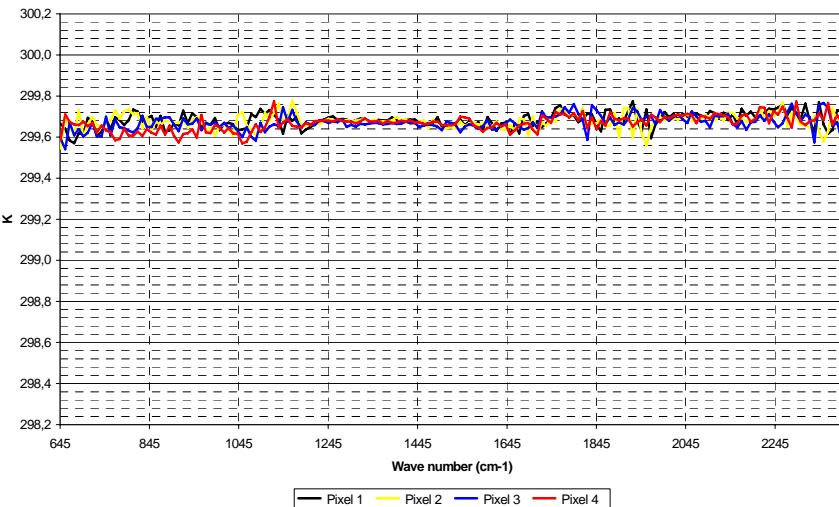




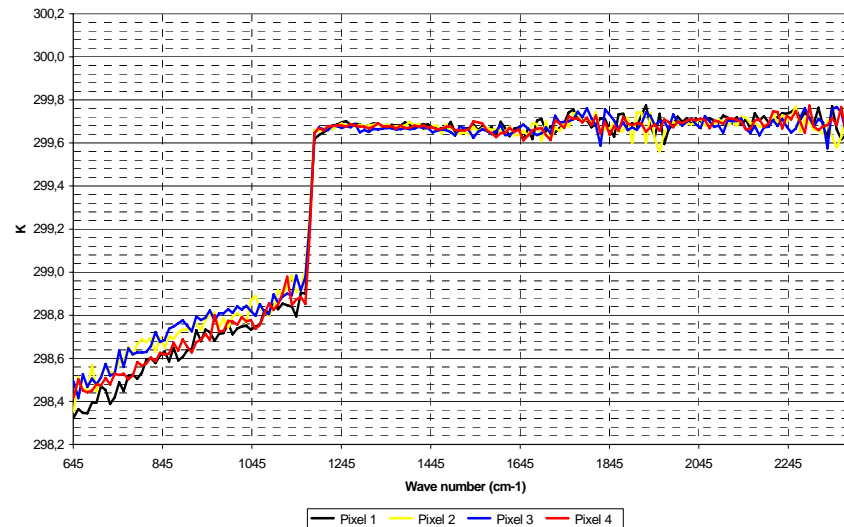
Non-Linearity correction efficiency

- Measurements performed with CBS temperature = 98,3 K
 - No NL correction : error = 1 K (B1)
 - NL correction : degree 2 polynom (B1)
 - Computed from interf at 98,2 K
→ error < 0.2 K
 - Computed from interfs at 93,5 K
→ error < 0.4 K

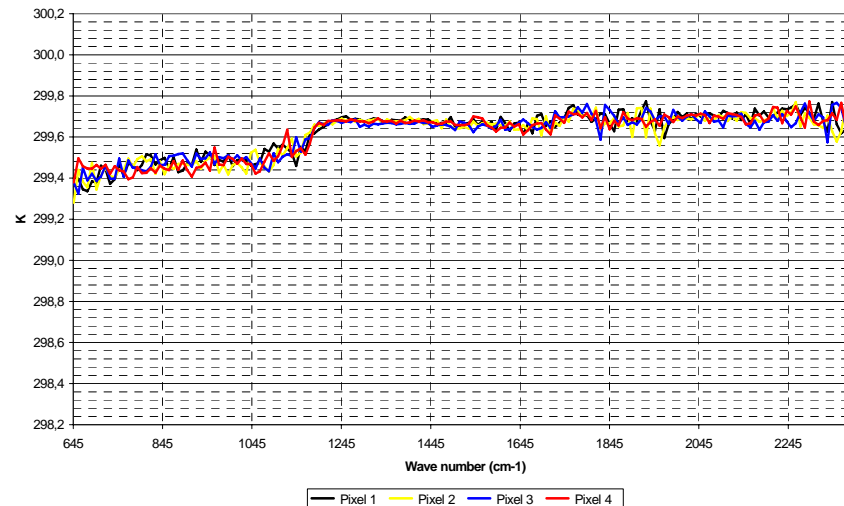
Session SM 06 22h59 — T_HBB about 299 K, CBS 98,3 K
Non-Linearity correction calculated on same session



Session SM 06 22h59 — T_HBB about 299 K, no Non-Linearity correction



Session SM 06 22h59 — T_HBB about 299 K, CBS 98,3 K
Non-Linearity correction calculated on session sm0306150454 (CBS 94 K)





Spectral measurements during FM2 tests

- An integrating sphere is put in front of the instrument with 2 laser beams injected simultaneously in it for direct measurement of the Instrument Spectral Response Function (ISRF)
 - In B1 band : laser CO2 (10.59 μm)
 - In B3 band : laser HFDF (3.76 μm)
 - Laser beams are periodically occulted by a shutter for measurement of the sphere background emission
 - Gives also a direct measurement of the Cube Corner trajectory
- Results are computed from raw interferograms processing
- Test performed with instrument either in
 - **External Calibration mode**
 - Allows to acquire a lot of data
 - Scan motion is limited to beta compensation law (biggest part of the dynamic disturbance)
 - **Normal Operation mode**
 - Nominal operation mode
 - But lower amount of data available
- Sounder IPSF measured with a dedicated OGSE (Geometric Collimator)



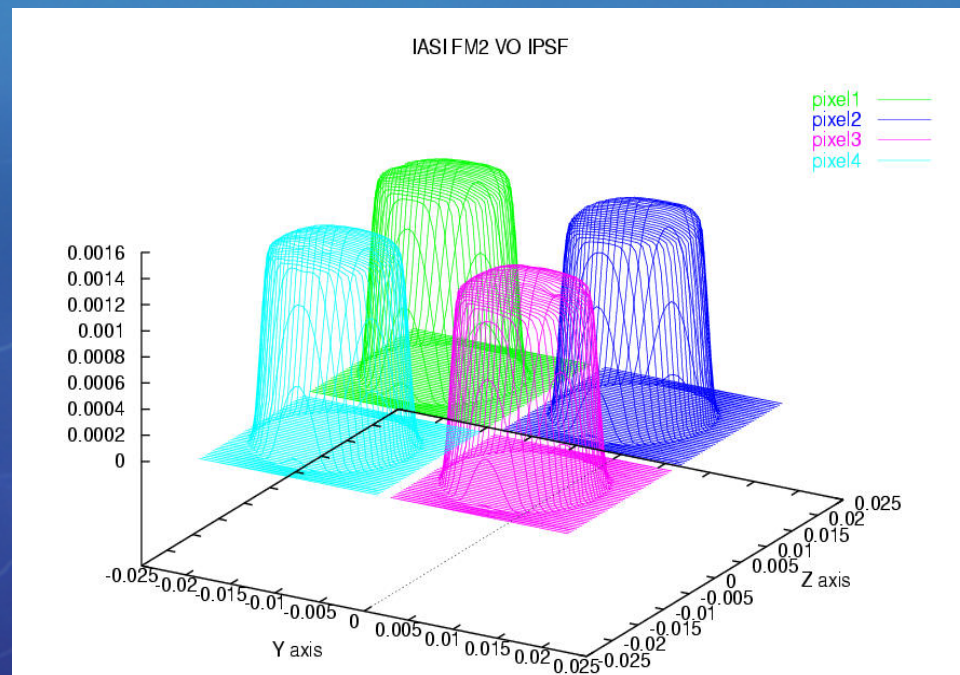


IPSF

Non conformity % peak to valley measurement

Specification : $\pm 5\%$

	B1	B2	B3
Pixel 3	3,7	3,8	3,7
Pixel 4	6,8	10,0	10,0
Pixel 1	4,6	2,3	3,8
Pixel 2	2,6	3,0	4,4



IPSF measurements close to predictions

Alignment of the interferometer :

axis position $< 200 \mu\text{rad}$

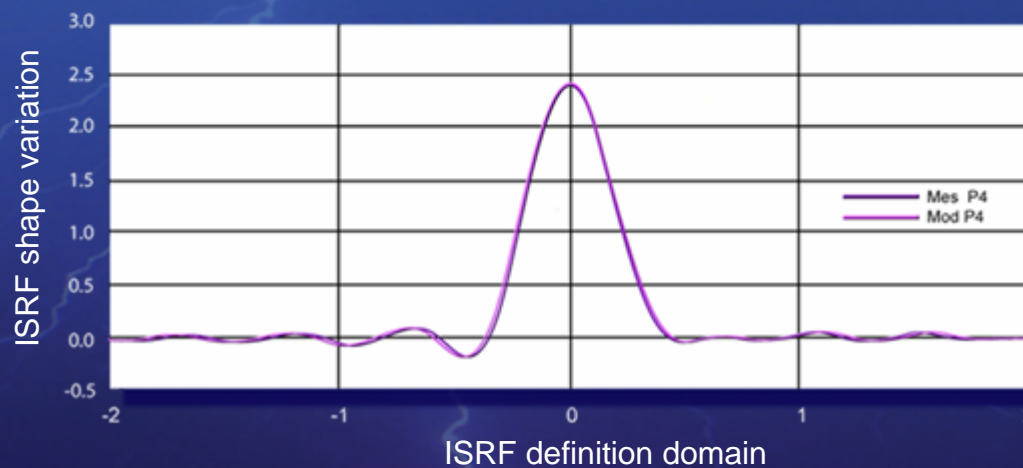
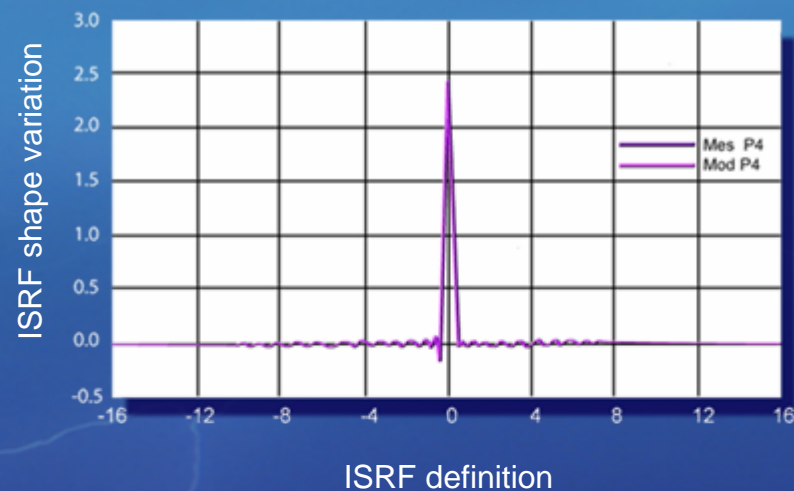
Cube corner offset : $32 \mu\text{m}$





Spectral

- The Instrument Spectral Resolution Function which was measured during the optical vacuum test is similar to the predicted one.





Requirements

Principle

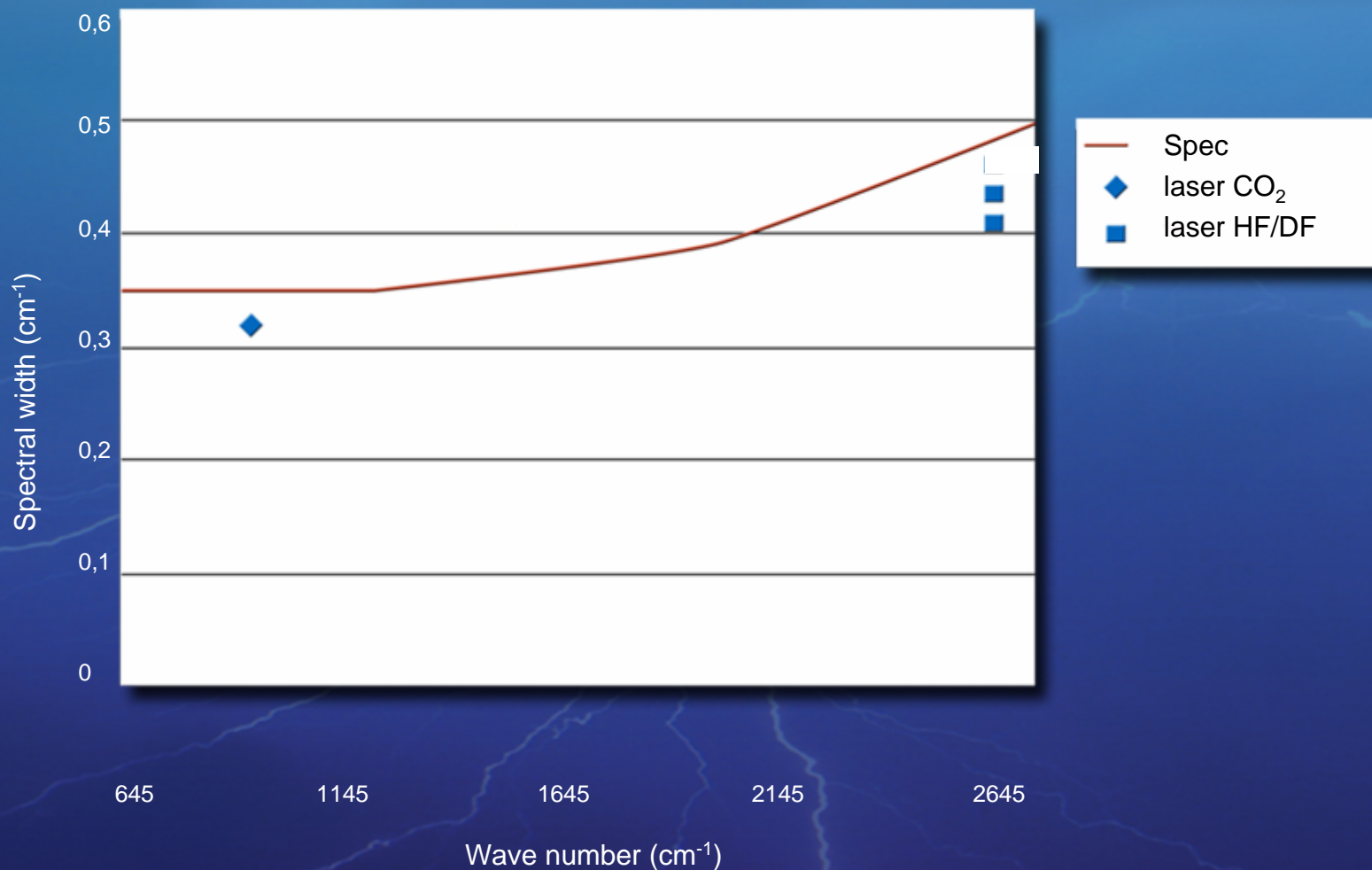
Design

Sub systems

Performances

Development plan

Spectral resolution



Spectral resolution within the specifications for the 4 pixels.



Spectral resolution

Maximum spectral shift = $10^{-5} < \text{spec. } (2 \cdot 10^{-4})$

Spectral stability : $< 8.3 \cdot 10^{-8}$ in B1
 $< 2.5 \cdot 10^{-8}$ in B3

Spectral stability well in the spec. ($1 \cdot 10^{-8}$)

Shape error index ε_2 : = quality index on the knowledge of the Instrument Line Shape = difference between modeled ILS and actual ILS

< 0.014 in B1 ~ spec. (0.026)

< 0.023 in B3 ~ spec. (0.042)

Within the specifications

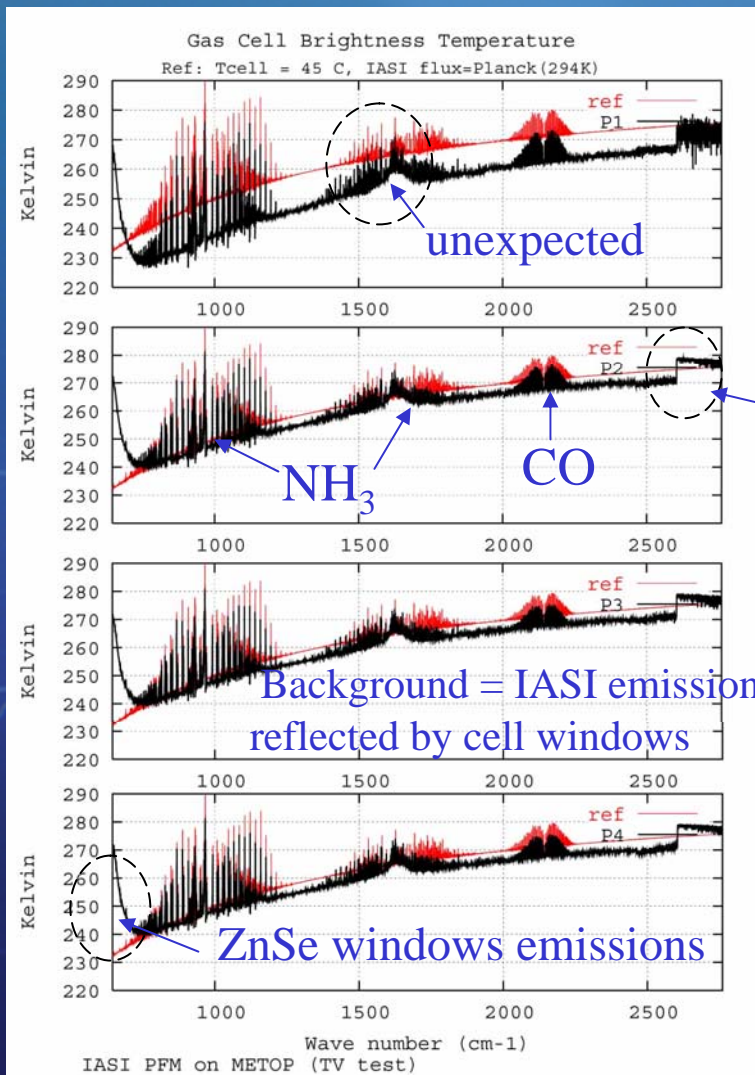
Shape error index ε_1 : = index on the knowledge of the ISRF

ILS knowledge meets the spec. marginally



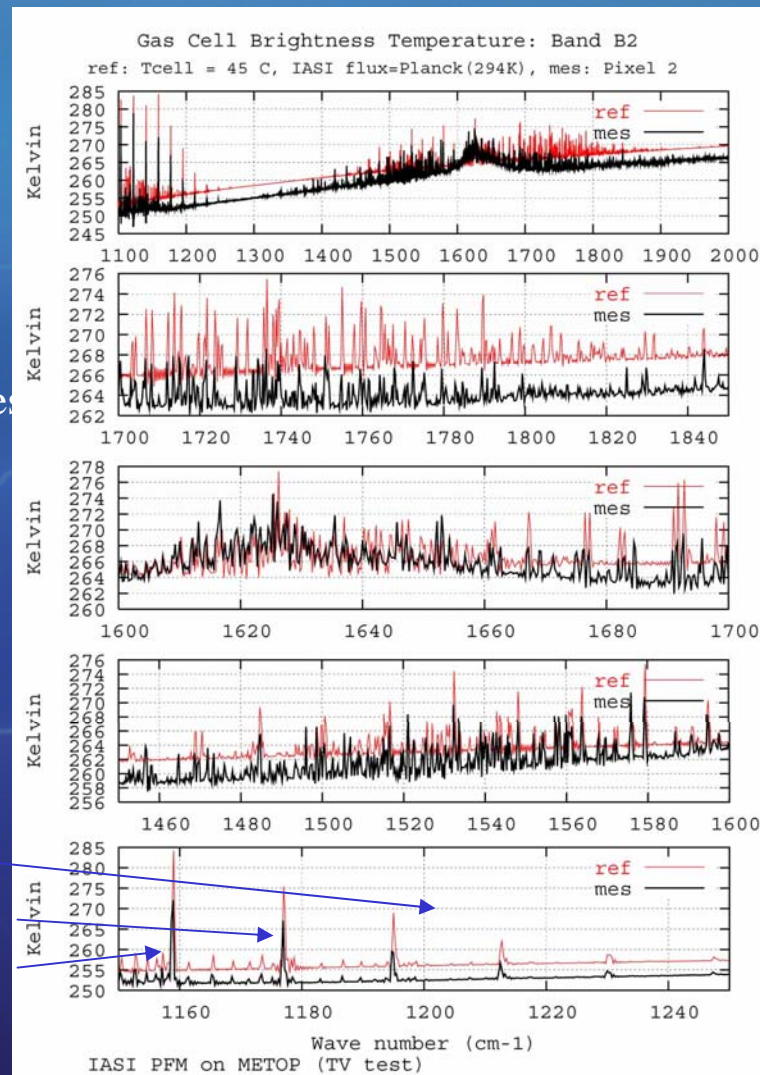


Gas Cell Measurements during TV test on Metop 2



"optimized" coding table

Spectral lines in B1/B2 interband





Geometry

- Pixels geometrical characterisation
- Pixel diameter : compliant with the expected values
- with less than 0.1 mrd difference
- Angular distance (in mrd) between pixel radiometric centers (nominal value is 21.63 mrd)

	P1P2	P2P3	P3P4	P4P1
B1	21,65	21,61	21,67	21,61
B2	21,52	21,68	21,57	21,55
B3	21,55	21,66	21,57	21,57

- Effective pixel centres included within circles of less than 0.15 mrd in radius centered on the reference square corners (specification is 2 mrd)

Geometric performances well within the spec.





Imager

- Radiometric performances
 - NEDT = 0.57 K (specified at 0.80 K).
 - Calibration accuracy < 1 K
 - Dynamic range : noise better than specification from 200 K to 300 K of scene temperature
- Geometrical performance
 - 4 blind pixels (2 % specified)





FM3 Optical vacuum tests

- FM3 OVT tests were carried out successfully at Alcatel in October 2005
- FM3 showed the same behavior than FM2
- Radiometric performances are slightly better (specially in B1) and confirm the validity of the models.
- Calibration is as good as for the FM2
- Water contamination is weak and could lead to a loss of about 0.2%/day
- Instrument was accepted by the Technical Review board
- It is now stored at Alcael for integration on a next MetOp platform

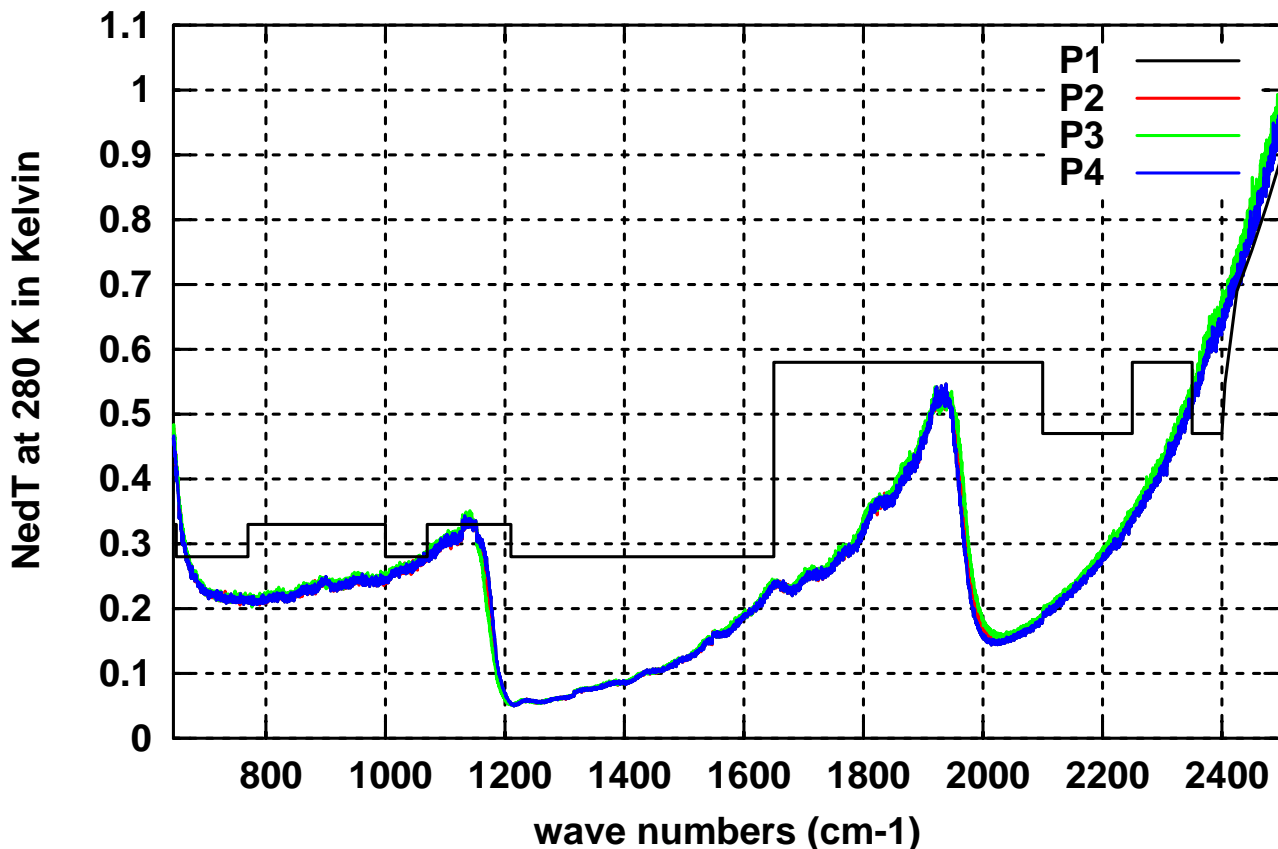




IASI FM3 on ground radiometric performances

sm0510241525

NOISE SPECTRUM LN [40..146]



31-10-05 14:32

IASI FM3 Optical Vacuum Test

CNES DCT/PO/EV





Comparison IASI/AIRS

	AIRS on AQUA	IASI on MetOp
Overpass time	13:30	9:30 AM
Spectral coverage	2378 channels 650 to 2670 cm^{-1} with gaps (1136-1216, 1614-2170)	8461 channels 645 to 2760 cm^{-1} continuous
Radiometric performances	See plot	See plot
Spectral Performances	From 0.4 to 2.5 cm^{-1}	From 0.32 to 0.46 cm^{-1}
Geometry	3*3 pixels of 15*15 km^2	2*2 pixels of 12*12 km^2



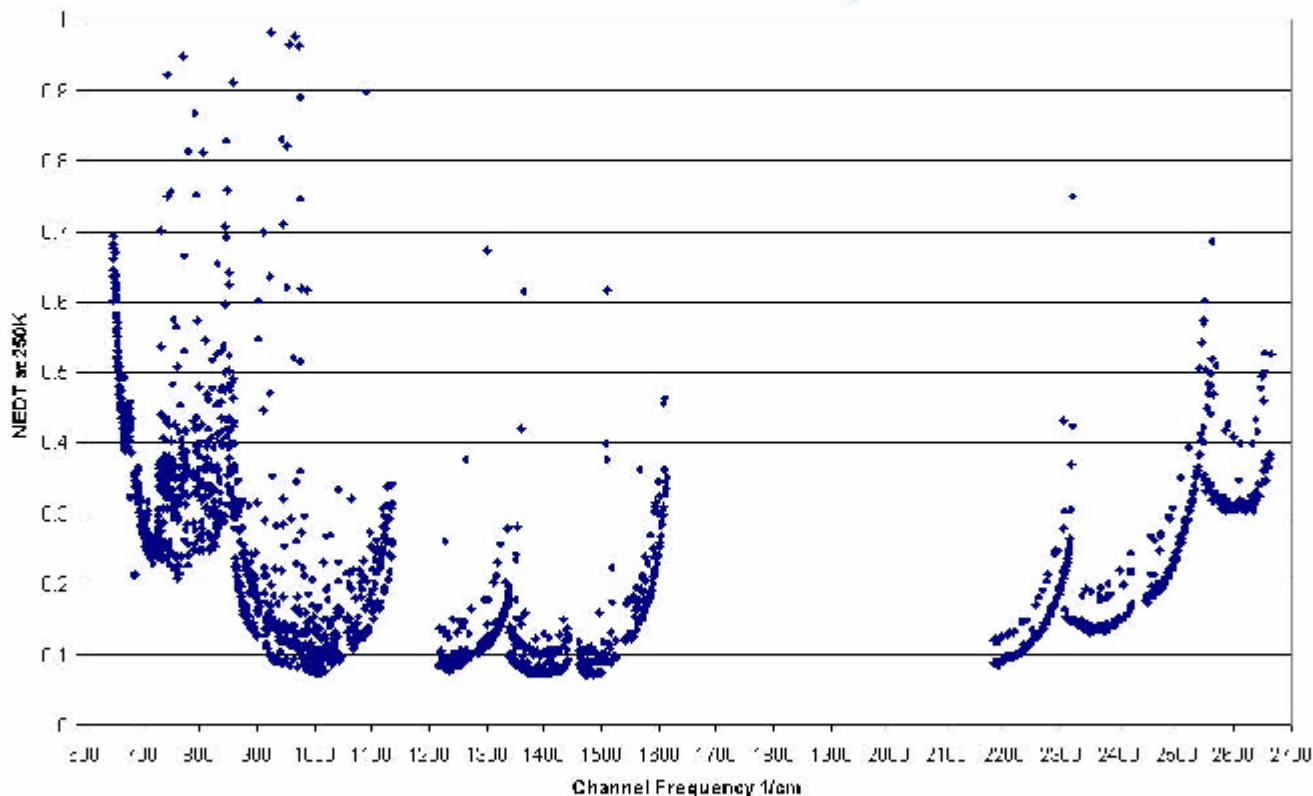


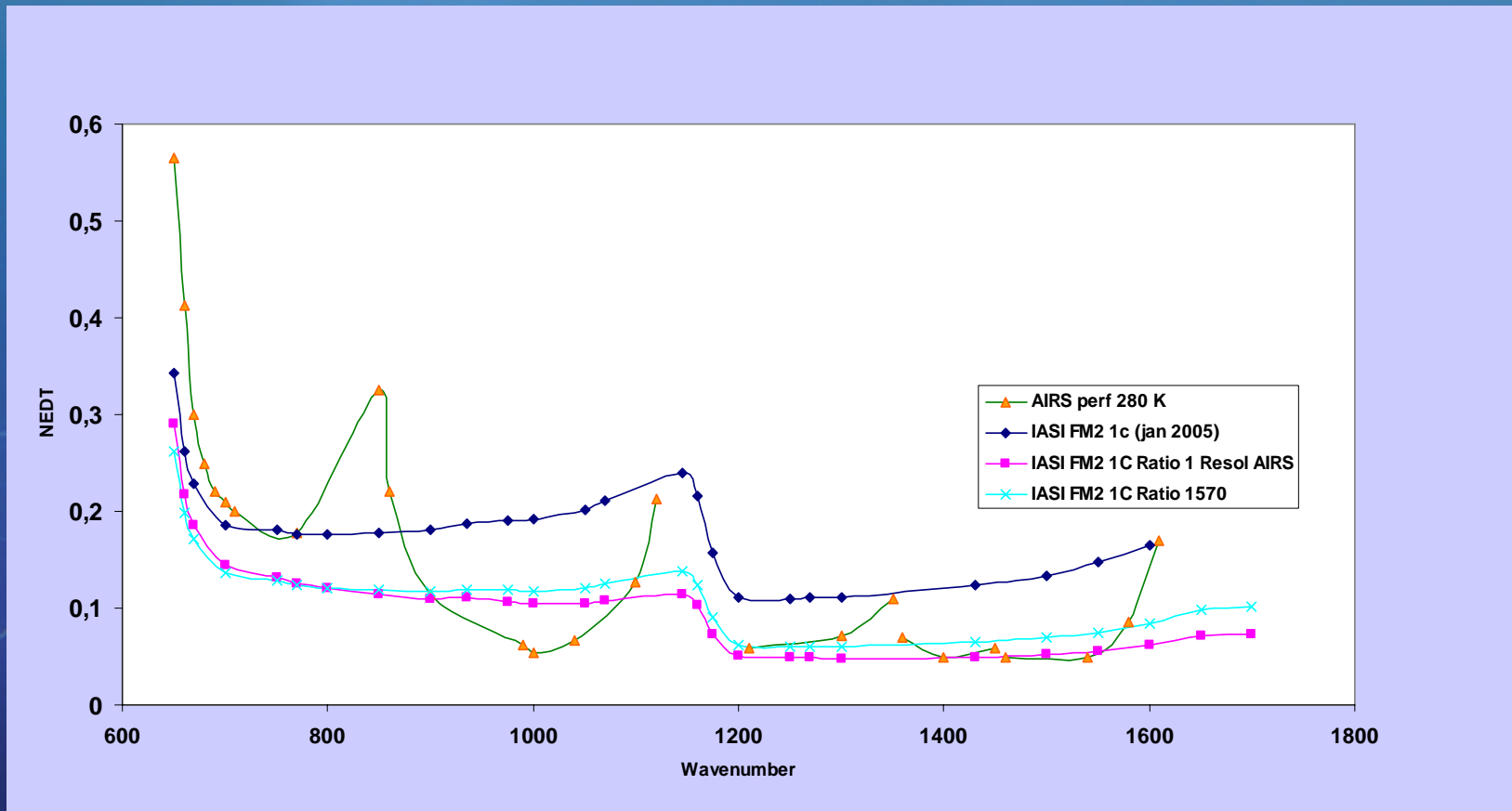
Measurement NEDT is excellent and stable

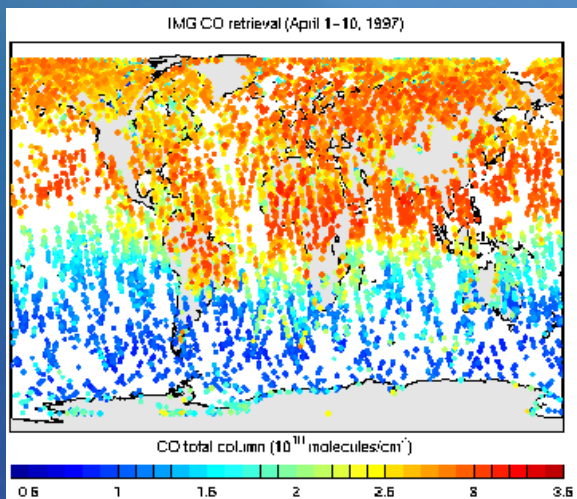


AIRS NEDT at 250K measured for 240 granules on 2 June 2003

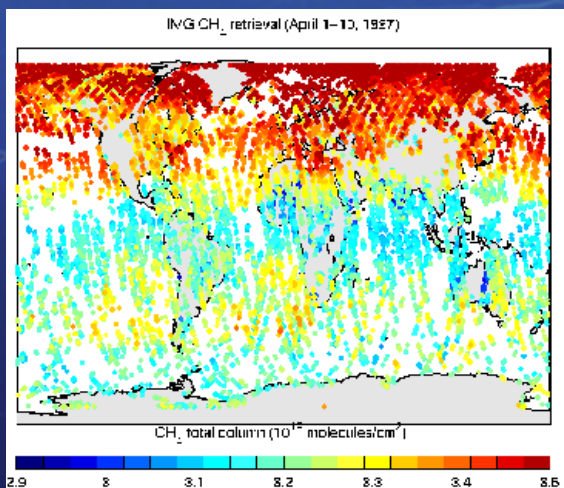
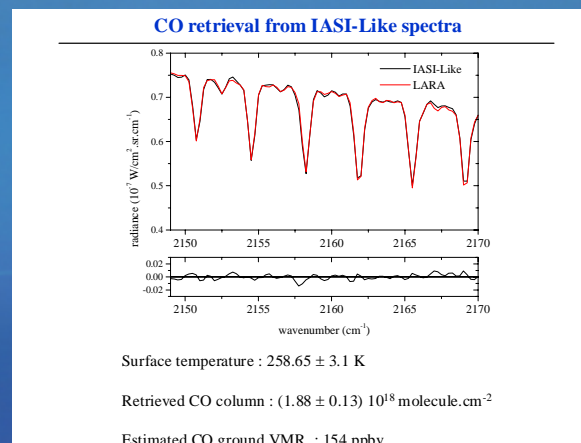
2187 of the 2378 AIRS channels have NEDT < 1K at 250K and gaussian noise characteristics



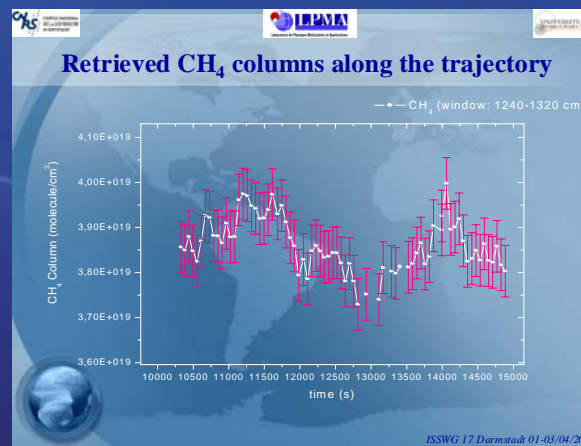


CO and CH₄

CO
column:
10-15%



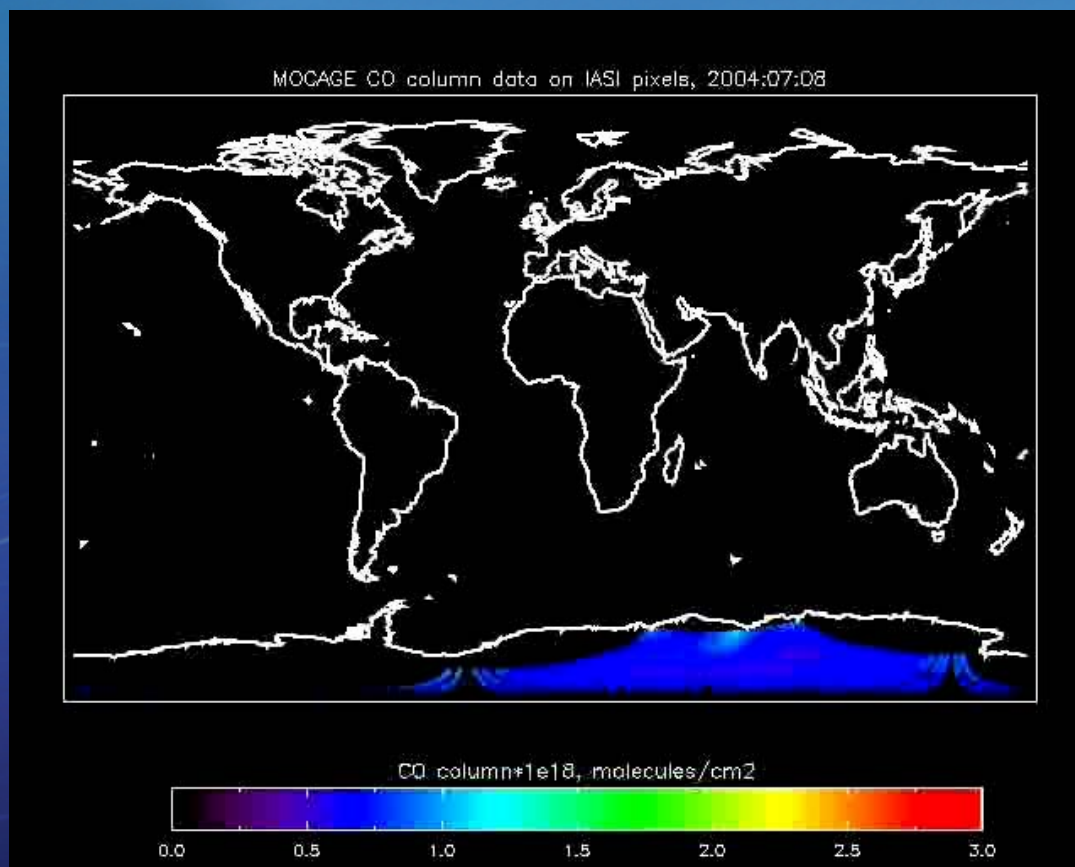
CH₄
column:
3%





One day CO with IASI in cloud free conditions

(courtesy C. Clerbaux)





Conclusions

- The last on the ground tests show that
 - All performances meet their specifications (sounder and imager)
 - IASI proved an excellent behavior
 - The performance model is good and can be used in the analyses
- IASI will be an tremendous step forward in meteorology and atmospheric research



Infrared Atmospheric Sounding Interferometer



THANK YOU FOR YOUR ATTENTION

More information on
<http://smc.cnes.fr/lasi>



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